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MOUNT PAKTU RESEARCH CENTRE



“Geomagnetic Variations and Space Weather Hazard Assessment in the Mongolian Region”

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- ✓ Research Objectives
- ✓ Physical Background
- ✓ Data
- ✓ Results
- ✓ Conclusion

Research Objectives

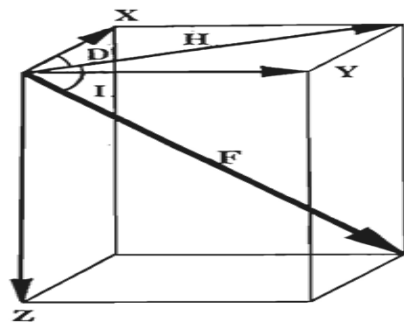
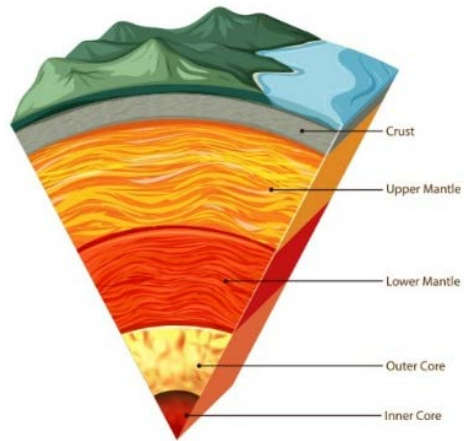
- Develop a long-term geomagnetic storm catalog for Mongolia covering the period 1995-2026 and examine its relationship with solar activity (Sunspot Number, SSN).
- Investigate the physical mechanisms of the 5-6 December 2020 sudden-commencement (SSC) CME-driven geomagnetic storm through integrated analysis of satellite and ground-based observations.
- Examine magnetopause compression, magnetic reconnection, and the propagation delay of the solar wind disturbance to improve understanding of Sun Earth interactions and regional geomagnetic responses.

Significance

- Enhance the regional geomagnetic storm database for space weather research in Mongolia.
- Demonstrate the application of integrated satellite and ground-based observations for investigating geomagnetic storm processes.
- Contribute to regional space weather hazard assessment through improved understanding of geomagnetic storm dynamics.

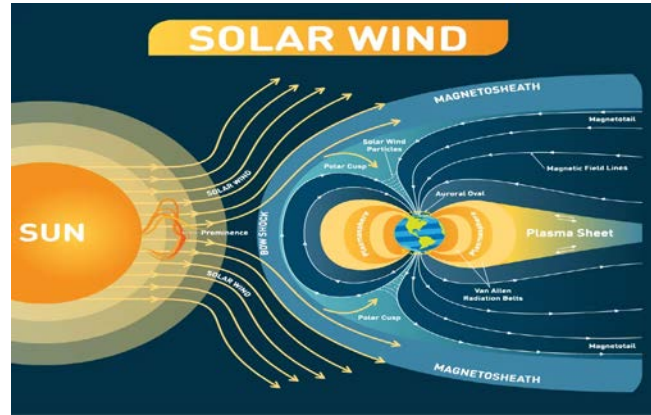
Solar Wind
↓
Magnetosphere
↓
*Ground
Observations*

Internal Magnetic Field Sources



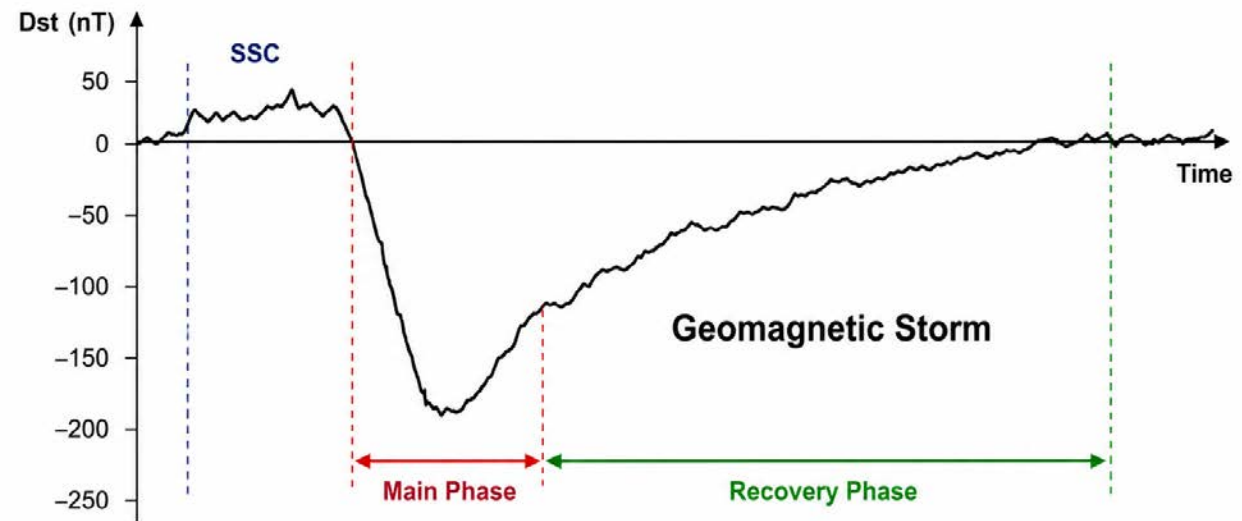
$$H = \sqrt{X^2 + Y^2}$$

External Magnetic Field sources

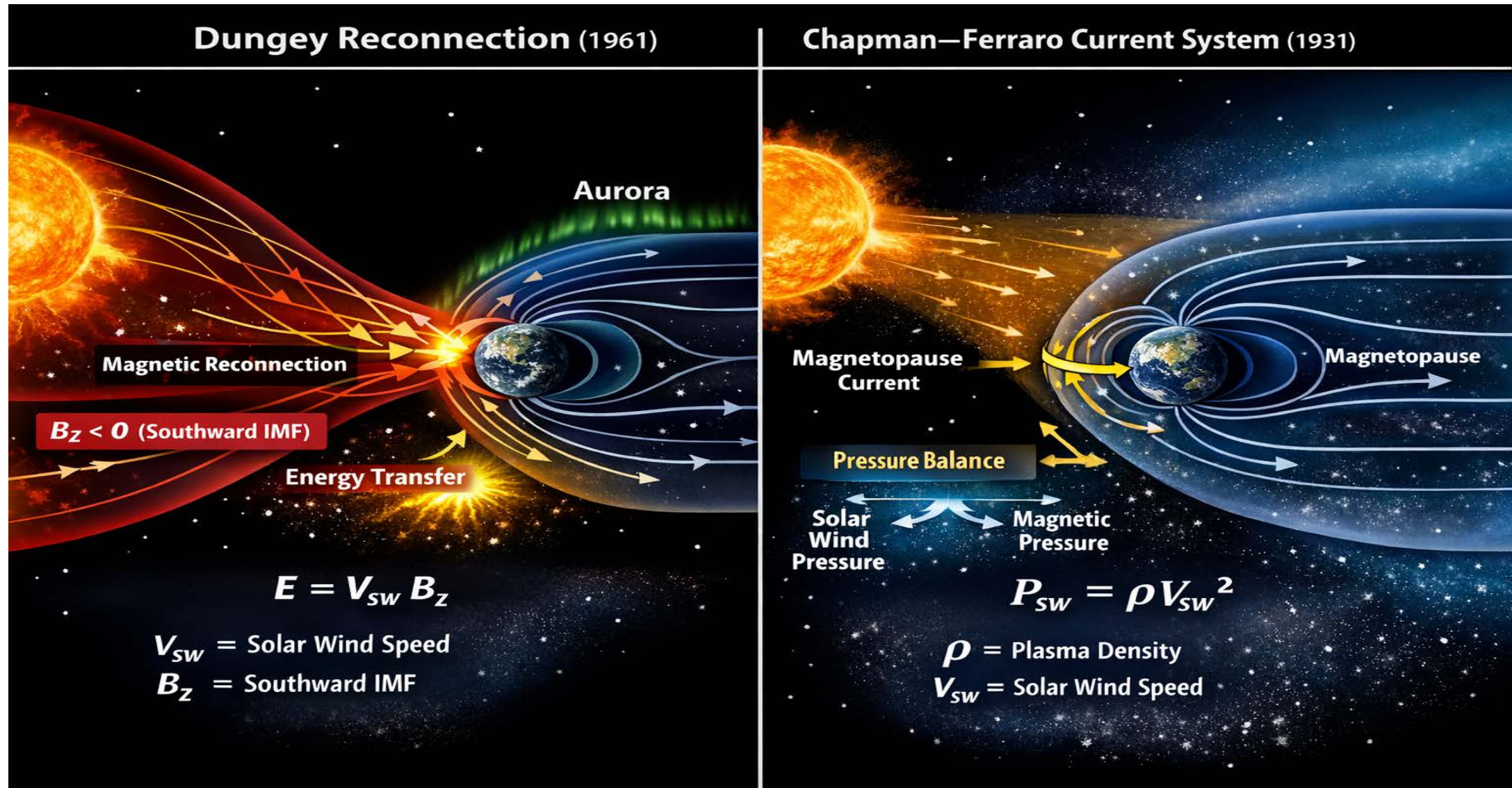


Magnetospheric Current Systems

- Magnetopause current
- Ring current
- Magnetotail current
- Field-aligned currents



Dst index indicates the intensity of geomagnetic storms



Magnetic reconnection and solar wind dynamic pressure are the primary physical drivers of geomagnetic storms.



OMNI Database

- Solar wind speed (V_{sw})
- B_z , IMF
- Dynamic pressure (P_{dyn})
- SYM/H index



Cluster Mission

- IMF (B_x , B_y , B_z)
- Plasma velocity (V_x)

Ground based magnetometer of Mongolia

HVD-MAGDAS

- (H , D , Z components)



INTERMAGNET

- Ground magnetic field
- H component



10 Hz magnetic
field data



1-minute averaging



Selection of 5
International Quiet
Days (IQD)



Baseline (H_q)

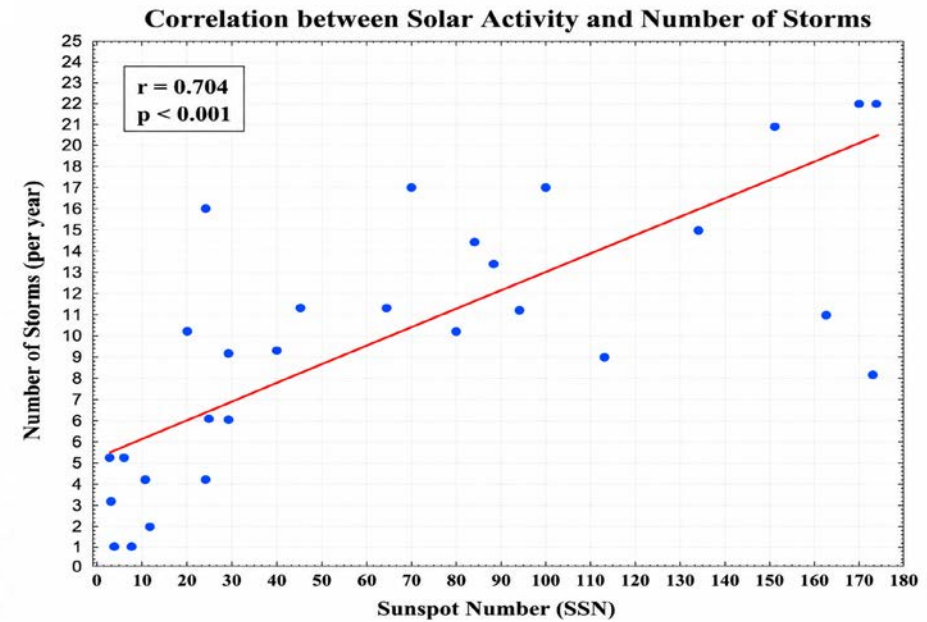
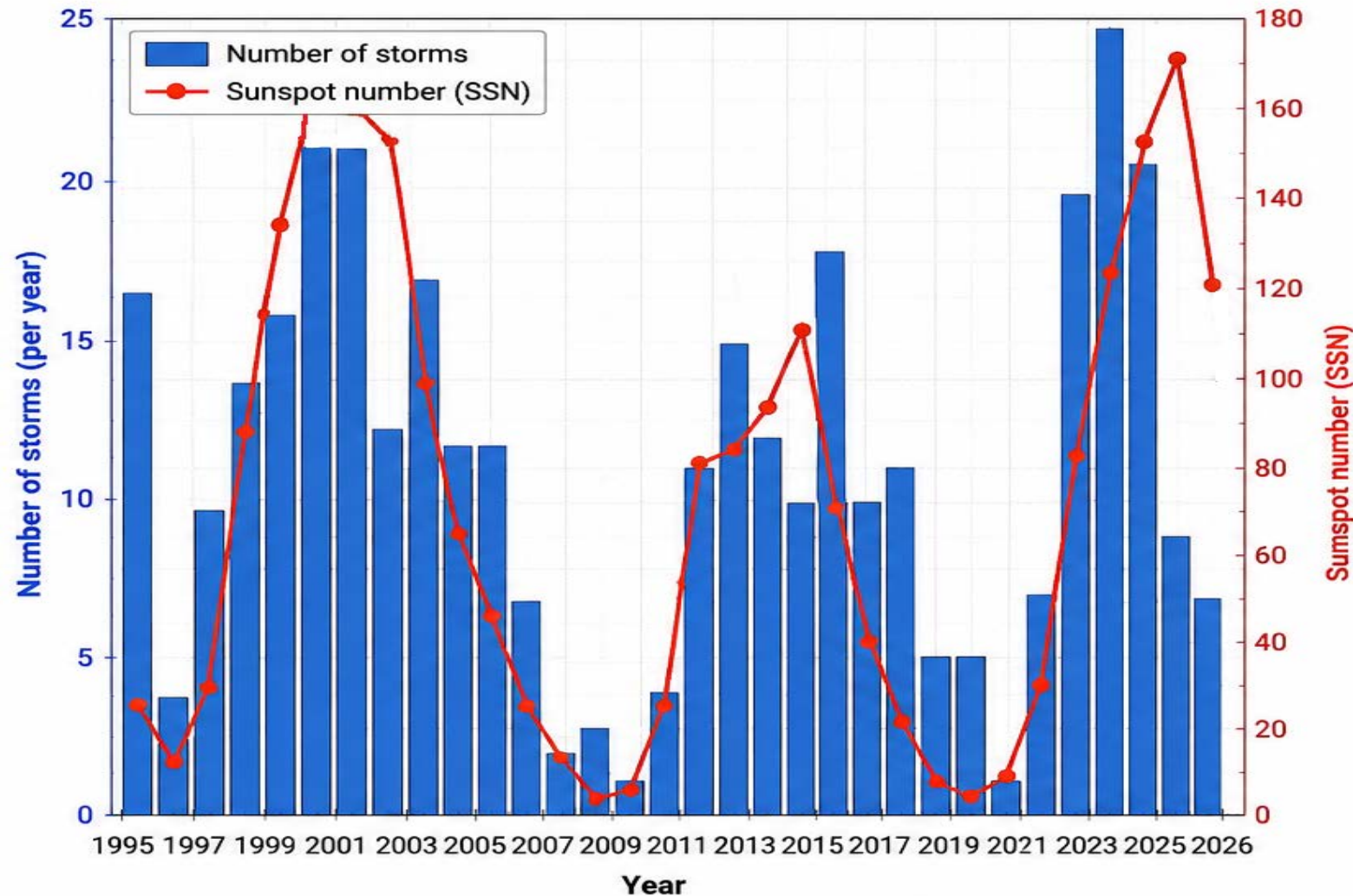


$\Delta H = H - H_q$

Results: Long term Statistics of Geomagnetic Storms (1995-2026)

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Geomagnetic Storms and Solar Cycle (1995–2026)

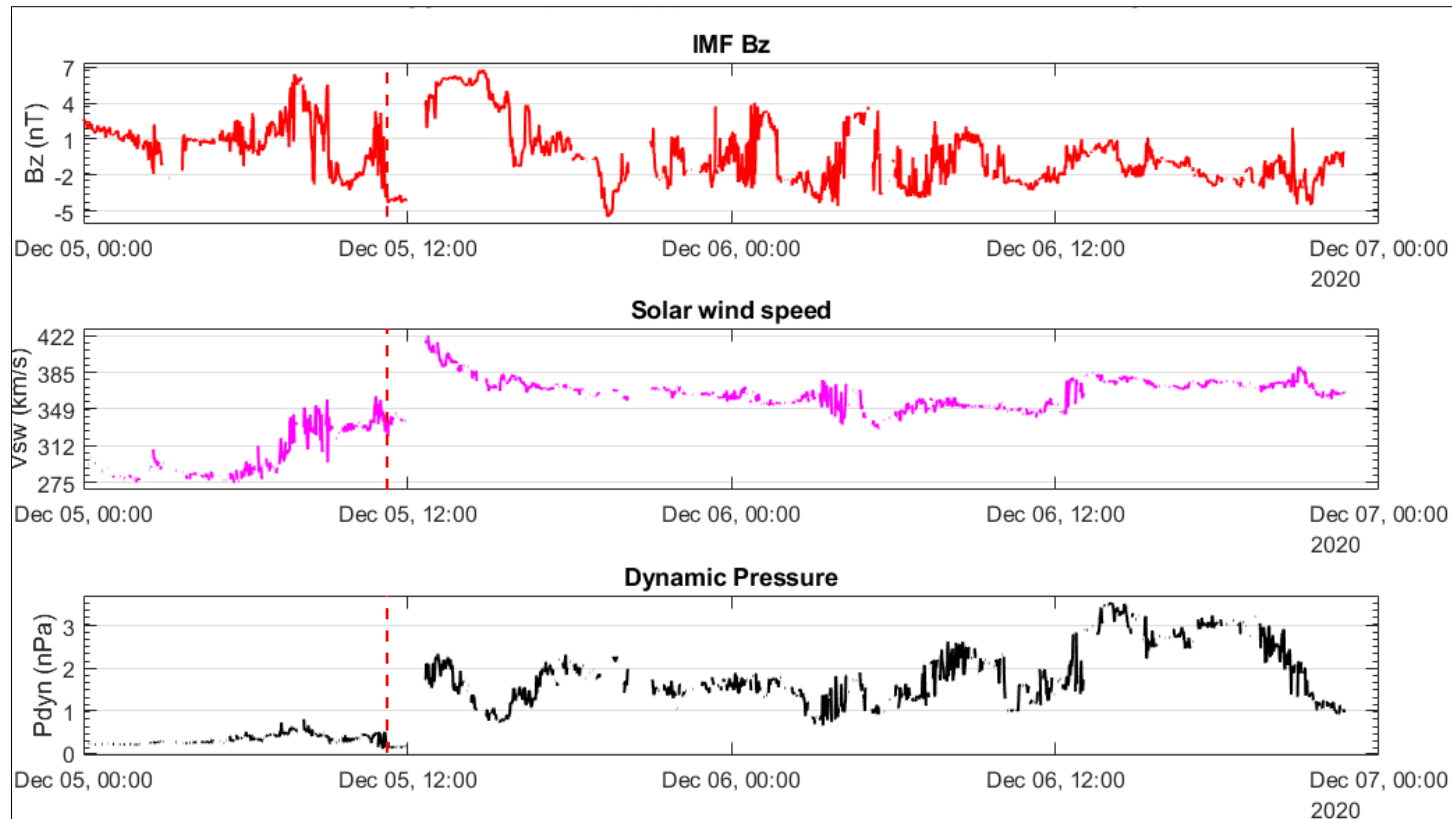


✓ **356** geomagnetic storms identified

✓ Strong positive correlation: $r = 0.704$

✓ Storm occurrence follows the solar cycle

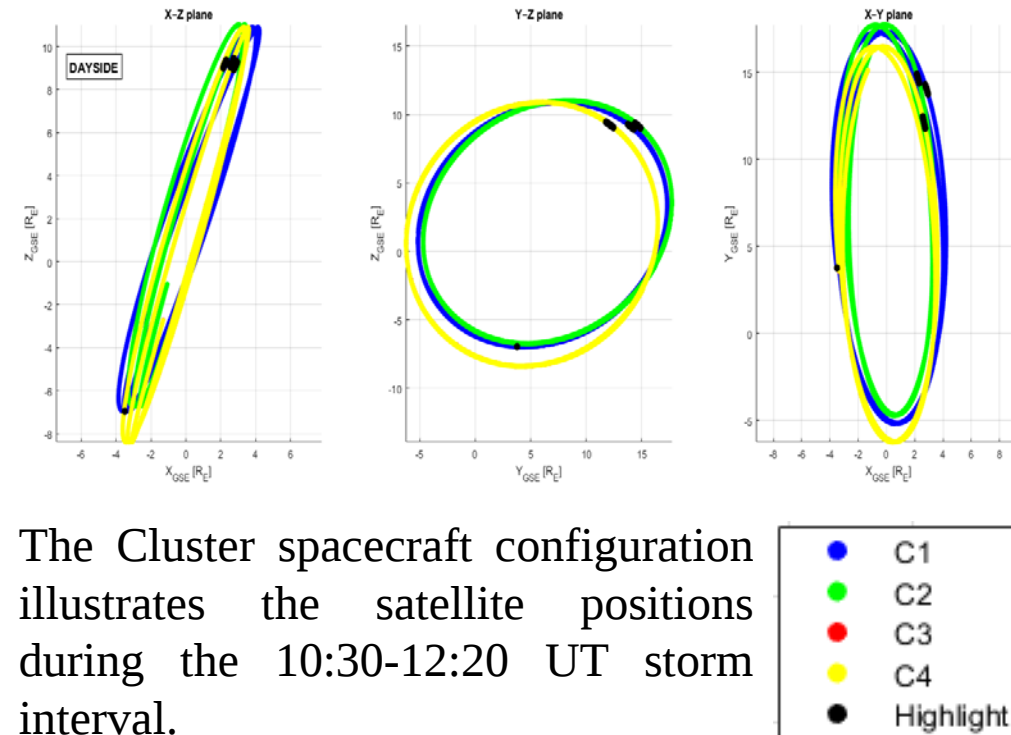
Magnetosphere



The OMNI observations show a southward turning of the IMF Bz, accompanied by increases in solar wind speed and dynamic pressure during the storm interval.

IMF Bz $\approx$ -4 nT
Vsw $\approx$ 422 km/s
Pdyn $\approx$ 2-4 nPa

Cluster Spacecraft Positions

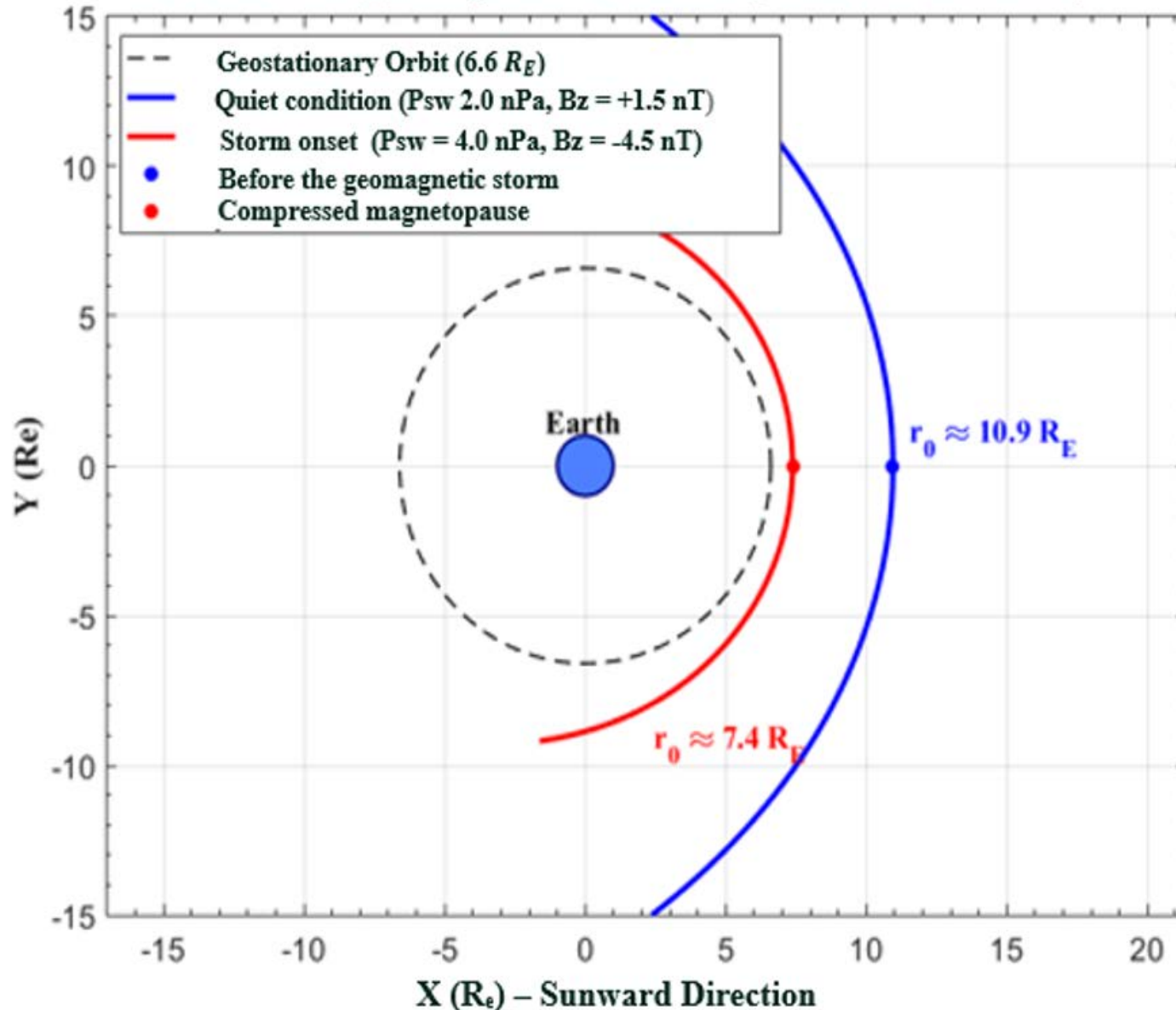


The Cluster spacecraft configuration illustrates the satellite positions during the 10:30-12:20 UT storm interval.

Case rationale

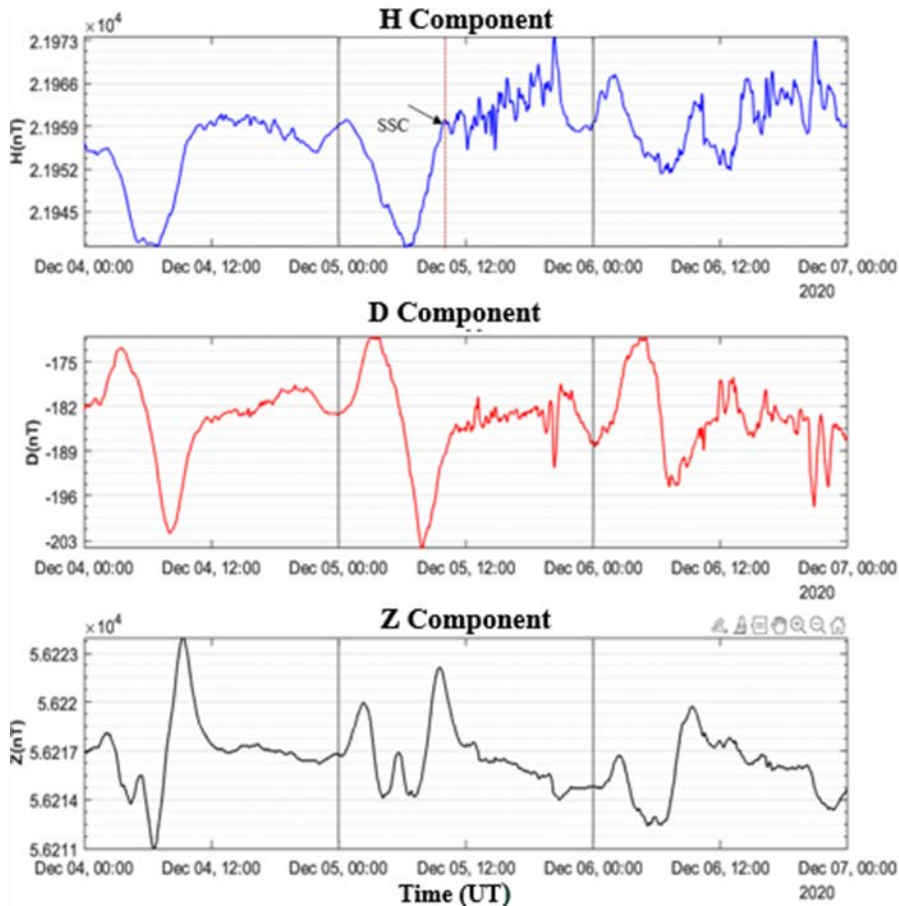
- Rare storm during low solar activity (Solar Cycle 24-25 transition).
- Clear regional response ($Dst < -40$ nT).
- Multi-source observations (Cluster+ground).

Modeled Magnetopause Position (Shue et al., 1998)



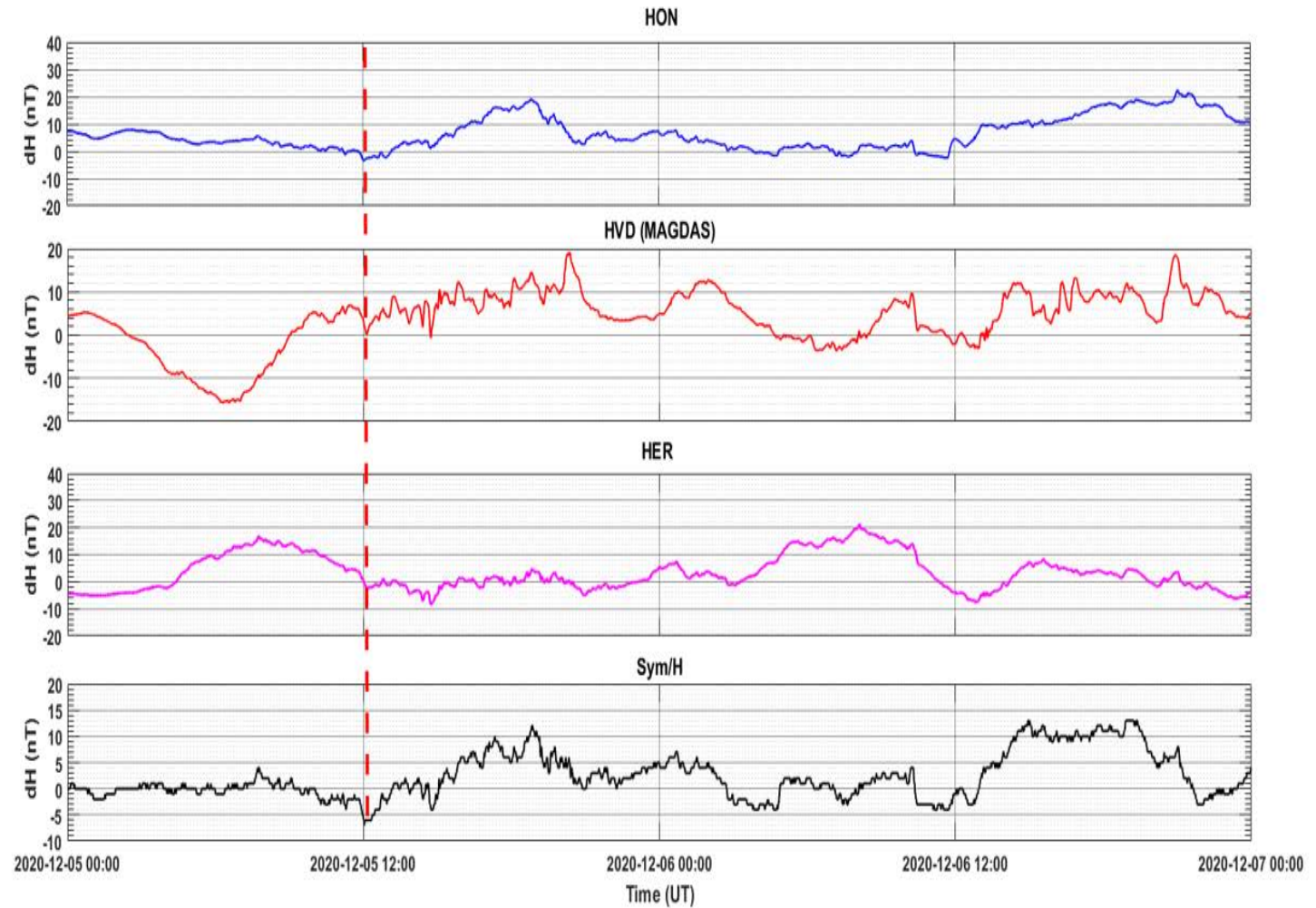
- Enhanced solar wind dynamic pressure compressed the dayside magnetopause.
- The subsolar stand-off distance decreased from $\sim 10.9 R_E$ to $\sim 7.4 R_E$ ($R_E = 6370 \text{ km}$)
- The modeled magnetopause compression is consistent with the observed geomagnetic sudden impulse (SI).

Hovd Observatory (HVD-MAGDAS) Magnetic Field Components (H, D, Z)



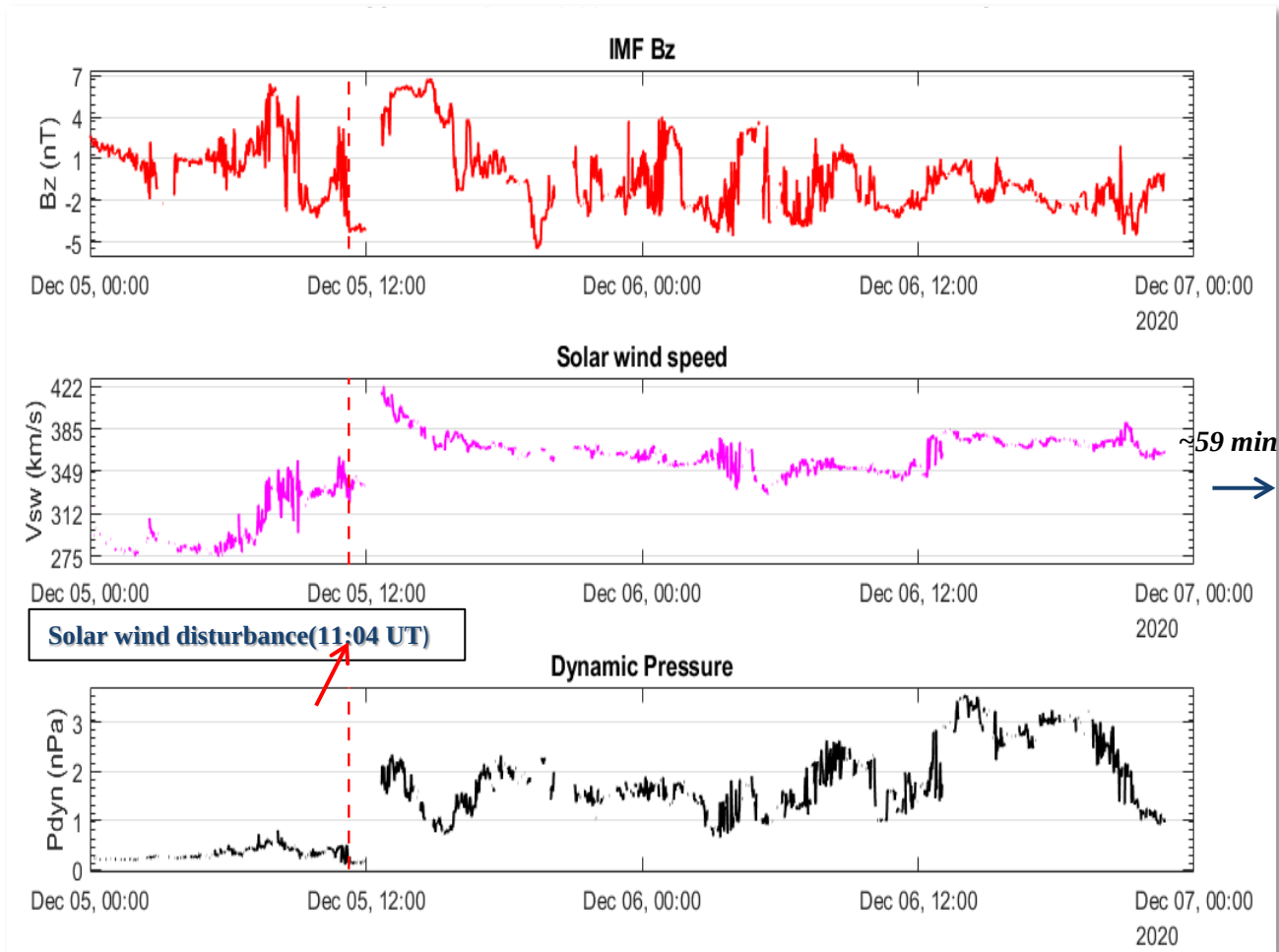
H
↓
 ΔH

ΔH - Variations of Geomagnetic field

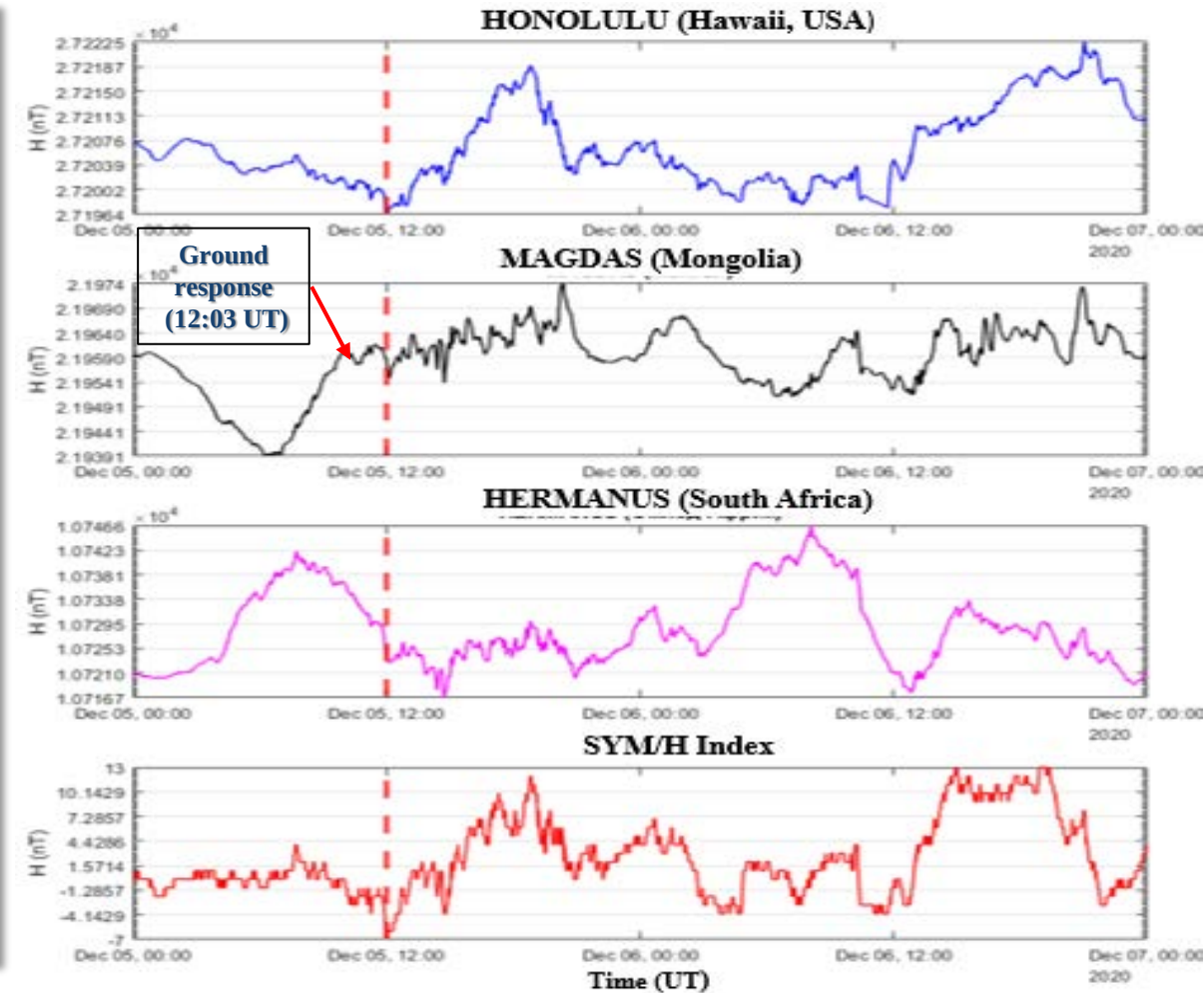


- SSC was identified in the H component at HVD in Mongolia
- Consistent ΔH responses were observed at HON, HVD, and HER.

Observations of Satellite (OMNI)



Observation of ground based Magnetometer



The solar wind disturbance detected at 11:04 UT was followed by a ground magnetic response at 12:03 UT (~59 min later), with the SYM-H index confirming the associated global geomagnetic disturbance.

In this study, geomagnetic variations and geomagnetic storms observed in Mongolia during the period 1995-2026 were investigated using geomagnetic data from the Ulaanbaatar Observatory. Based on these observations and solar wind data, the physical conditions responsible for space weather hazards, the response of the Earth's magnetic field, and the potential for early warning were examined.

- ❖ The southward component of the interplanetary magnetic field ($B_z < 0$), together with sudden increases in solar wind dynamic pressure (P_{sw}) and solar wind velocity (V_{sw}), are the primary physical drivers of geomagnetic storms. These variations produce significant disturbances in the Earth's geomagnetic field and play a key role in the development of hazardous space weather conditions.
- ❖ The response of the Earth's geomagnetic field to changes in the solar wind parameters was found to occur with an average time delay of 49–59 minutes. This time lag indicates that geomagnetic storms can be forecast approximately 50 minutes in advance using real-time solar wind observations.
- ❖ The findings of this study provide a scientific basis for the development of space weather early warning systems aimed at protecting critical infrastructure, including satellites, electric power transmission networks, communication systems, and navigation services.

The results therefore have significant scientific and practical value for space weather monitoring and hazard mitigation. This study demonstrates that satellite and ground-based observations provide a consistent framework for investigating geomagnetic storm processes and assessing regional space weather hazards.

Thank you for your attention